

**Physics 530: Fall 2004**

**Homework 1**

**Due: Monday, Sept. 13th**

1. (i) Given that  $T^\mu{}_{\nu\rho}$  is a tensor of type  $(1, 2)$ , show that  $S_\mu = T^\rho{}_{\mu\rho}$  is a covariant vector.  
(ii) Show that if  $A$  is a 2-form then  $dA$  is a 3-form, that is, verify that  $dA$  is a tensor. (Do this by explicitly checking the transformation laws.)  
(iii) Show that if  $A$  is a 2-form then  $d(dA) \equiv 0$ .

*Comment: Parts (i) and (ii) are, of course, true for an arbitrary  $k$ -form. I am simply trying to save you some writing by asking you to check it for a 2-form. In doing this exercise you should note that there is nothing special about  $k = 2$ .*

2. Consider the unit sphere in Cartesian coordinates in  $\mathbb{R}^3$ :

$$S^2 \equiv \left\{ (x^1, x^2, x^3) : (x^1)^2 + (x^2)^2 + (x^3)^2 = 1 \right\}. \quad (1)$$

In lectures I introduced the stereographic coordinate system defined on  $S^2 \setminus N$ , where  $N$  is the North pole (at  $x^3 = 1$ ):

$$(y^1, y^2) \equiv \left( \frac{2x^1}{1-x^3}, \frac{2x^2}{1-x^3} \right). \quad (2)$$

There is also the natural chart of polar coordinates in which

$$(x^1, x^2, x^3) = (\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta), \quad (3)$$

for  $0 < \theta < \pi$ ,  $-\pi < \phi < \pi$ .

- (i) Compute, and simplify  $y^\mu(\theta, \phi)$ .

*Comment: You might find it easier to parametrize everything using  $\frac{\theta}{2}$  rather than  $\theta$ .*

- (ii) Suppose  $W_\mu = (0, 1)$  is a covariant vector in the  $(y^1, y^2)$  system. What are the components in the  $(\theta, \phi)$  system? Suppose  $V^\mu = (0, 1)$  is a contravariant vector in the  $(y^1, y^2)$  system. What are the components in the  $(\theta, \phi)$  system?

- (iii) Invert the coordinate change (2) so as to obtain the Cartesian coordinates,  $(x^1, x^2, x^3)$ , of a point on the sphere in terms of  $(y^1, y^2)$ . This expression defines a map (an embedding) of  $S^2$  into  $\mathbb{R}^3$ . Consider the constant covariant vector field  $U_a = (1, 0, 0)$  on  $\mathbb{R}^3$ . Compute its pull-back,  $U_\mu$ , to  $S^2$  in the  $(y^1, y^2)$  system. Where does  $U_\mu$  vanish identically? Why does  $U_\mu$  vanish at some points given that

$U_a = (1, 0, 0)$  never vanishes?

(iv) Consider a map  $\phi : S^2 \rightarrow S^2$  in which a point  $(u, v)$  in the  $(y^1, y^2)$  chart is mapped to another point in the same chart via:

$$\phi : S^2 \rightarrow S^2 \quad \text{where} \quad \phi : (u, v) \rightarrow \hat{\rho}^{-2}(2u, 2(\rho^2 - 1)), \quad (4)$$

with

$$\rho^2 \equiv \frac{1}{4}(u^2 + v^2), \quad \hat{\rho}^2 \equiv \frac{1}{4}(u^2 + (v + 2)^2). \quad (5)$$

In particular, this map takes the great circle at  $y^1 = 0$  into itself. I claim that the field,  $U_\mu$ , that you computed in part (iii) is *invariant* under transportation via  $\phi$ , (i.e.  $\phi$  is a symmetry of  $U_\mu$ ). The proof in general is an algebraic mess, and so to simplify things **show that on the the great circle at  $y^1 = 0$** , the map,  $\phi$  transports the field  $U_\mu$  into itself.

*Comment: This symmetry is far from obvious in the  $(y^1, y^2)$  system. Can you figure out why  $U_\mu$  is obviously symmetric under this map? Think about where you got  $U_\mu$  from in the first place .... (This is not a required part of the answer to this question.)*

3. The standard Euclidean metric in Cartesian coordinates in  $\mathbb{R}^3$  is given by:

$$ds^2 = (dx^1)^2 + (dx^2)^2 + (dx^3)^2. \quad (6)$$

(i) Consider new coordinates  $(r, \theta, \phi)$ , where:

$$(x^1, x^2, x^3) = (r \sin \theta \cos \phi, r \sin \theta \sin \phi, r \cos \theta). \quad (7)$$

Compute the metric in these new coordinates. (Show your working!)

(ii) From your result in part (i), one can infer that the induced metric on the 2-sphere is:

$$ds^2 = d\theta^2 + \sin^2 \theta d\phi^2. \quad (8)$$

Now compute the metric on  $S^2$  in the stereographic coordinates (2). You should find that the metric takes what is known as the “conformally flat” form:

$$ds^2 = \Omega^2 ((dy^1)^2 + (dy^2)^2), \quad (9)$$

for some function,  $\Omega(y)$ . What is  $\Omega$ ?

4. (i) Let  $A$  be a  $k$ -form. Obtain a simplified expression (not involving  $\varepsilon$ -symbols) for the  $(k-1)$ -form  $\delta A \equiv \pm(-1)^{kn+n+1} * d * A$ . (Hint: Write  $\varepsilon$ -tensors in terms of  $\sqrt{|g|}$  and the  $\tilde{\varepsilon}$ -tensor densities).
- (ii) Write down a simple expression for  $\delta df$  where  $f$  is a function.
- (iii) Compute your expression in part (ii) for  $\mathbb{R}^3$ , but in the spherical polar coordinates (7). That is, use the metric from **3.** part (i).
5. Consider two points  $P$  and  $P'$  in a Newtonian gravitational field whose potential is  $\phi$ , and let  $\Delta\phi = \phi(P') - \phi(P)$ . Suppose one has an excited atom capable of emitting a photon of frequency  $\nu$ . Consider the following two processes.
- A. The atom emits a photon of frequency  $\nu$  at  $P$ , the photon passes through  $P'$  where its frequency is measured to be  $\nu'$ .
- B. The atom is carried to  $P'$  where it emits the photon of frequency  $\nu$ , and then the unexcited atom is carried back to  $P$ .

By calculating the work done compared to the photon energy measured at  $P'$  in both processes, show that

$$\frac{\nu'}{\nu} = (1 - \Delta\phi). \quad (10)$$

(Remember:  $E = mc^2$ ,  $E = h\nu$ .)